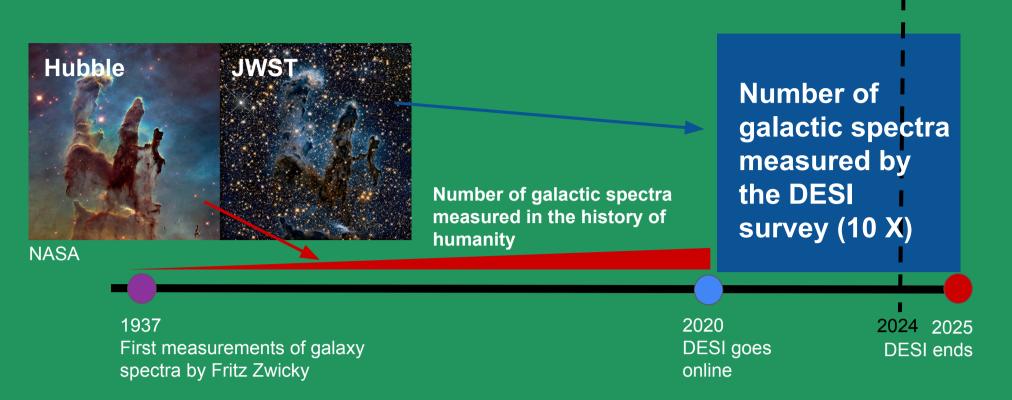


The problem: Cosmology is getting a lot of data



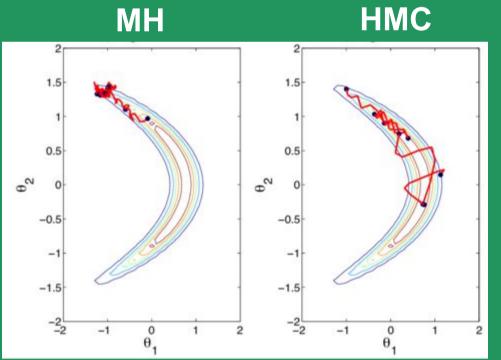
MORE **DATA**

MORE **PARAMETERS**

THAT WE **CANNOT CONSTRAIN**

- Traditional inference methods such as Metropolis hastings (MH) can constrain **tens** of parameters
- Up coming cosmological analysis will have hundreds of parameters to account for extra systematic effects

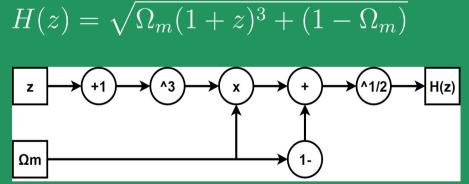
The Solution: Hamiltonian Monte Carlo (HMC)



Comparison between MH and HMC exploring a 2D parameter space by Lagrangian Dynamical Monte Carlo (Lan et al, 2012)

- HMC simulates hamiltonian trajectories to explore large parameter spaces very efficiently
- In order to do so, HMC requires the gradient of our theory predictions with respect to said parameters

How to get cheap gradients: Auto-differentiation



Symbolic representation of a computer program generated by auto-differentiation



Turing.jl

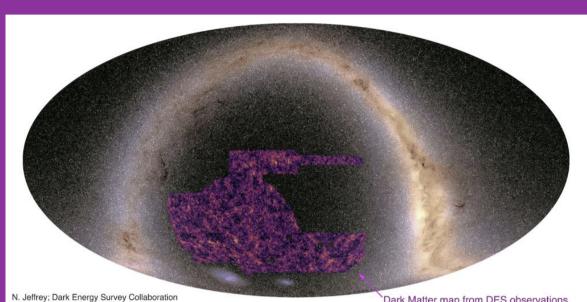


- Auto-differentiation allows computers to chain rule through computer programs by building a symbolic representation of the code.
- The Julia ecosystem Turing.jl: provides an statistical inference framework with a native HMC implementation
 - LimberJack.jl: provides auto-differentiable theory predictions for cosmological observables.
 - Combining the two high-dimensional cosmological analyses are possible

An Example: Auto-differentiable analysis of angular correlation functions

The Data:

the DES-Y1 data is composed of 300 million galaxies across 500 square degrees on the sky



Footprint of the DES Y1 survey compared to our view of the milky way

The Theory:

 $C_{\ell}^{UV} = \int \frac{d\chi}{\chi^2} q_U(\chi) q_V(\chi) P_{UV} \left(k = \frac{\ell + 1/2}{\chi}, z(\chi) \right)$

We predict the angular (on the sky) 2 point correlation function of these galaxies and their properties.

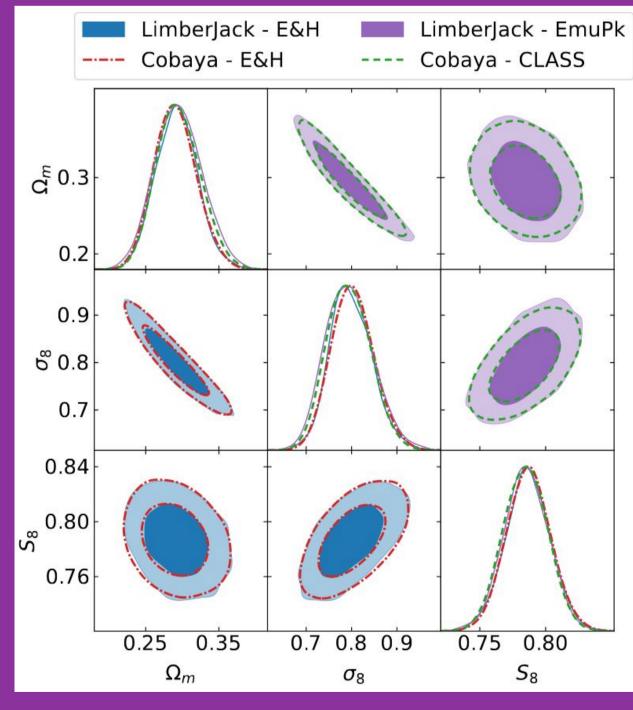
- LimberJack uses mix of emulation and auto-differentiable programming to evaluate these correlation functions and their gradients efficiently
- Emulation employs machine learning to build approximations to difficult computations

--- $\partial_{\Omega_k}C_l$ ⇔imberJack.jl --- $\partial_h C_l$ - $\partial_{\sigma_8}C_l$ --- $\partial_{n_s}C_l$ √10⁻⁸ Weak lensing -6 CMB lensing

Gradients of the different 2-point correlation functions with respect the five main cosmological parameters computed by LimberJack

Results:

- Identical parameter constraints than state-of art-software (Cobaya)
- Order of magnitude faster gradients than finite differences
- Enables efficient high-dimensional Cosmological analyses
- Gradient-based inference methods not worth it for low dimensions



Lower triangle compares LimberJack to Cobaya when both use approximations to compute their theory. **Upper triangle** shows comparisons when Limberjack employs

emulation to match exact computations. In this plot Cobaya uses MH and LimberJack HMC.



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Based of LimberJack.jl: auto-differentiable methods for angular power spectra analyses by J. Ruiz-Zapatero, D. Alonso, C. García-García, A. Nicola, A. Mootoovaloo, J. M. Sullivan, M. Bonici, P. G. Ferreira



